

# Dispersion in the Offset in fractal Dimension of Large Scale Structures

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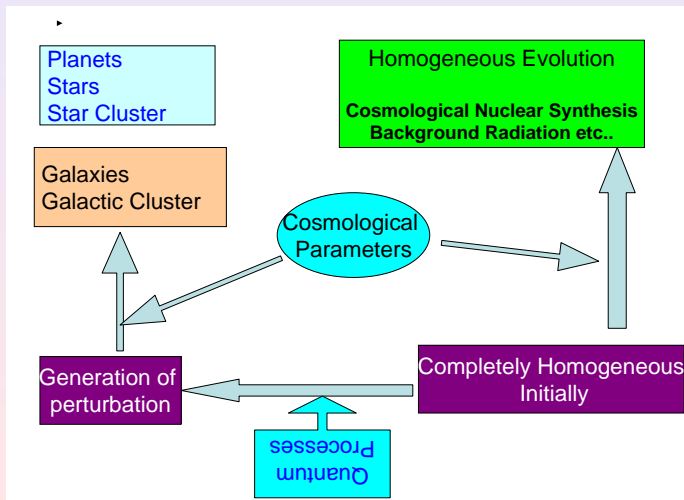
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# Pictorial Summary : Evolution of Universe



# Outline

- 1 Large Scale Structures (LSS)
  - Properties
  - Distribution in the Sky

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  - Issues Involved and Aims
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  - Fractal Dimension
- 3 Offset in fractal Dimension and the corresponding Dispersion
  - Relation of fractal Dimension to Correlation Function
  - Dispersion in the Offset

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- 3 Offset in fractal Dimension and the corresponding Dispersion
  - Relation of fractal Dimension to Correlation Function
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- 4 Results
  - Conclusions regarding Scale of Homogeneity
  - Finding the *real* scale of Homogeneity

# LSS of Universe

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=  $3 * 10^{18}$ cm

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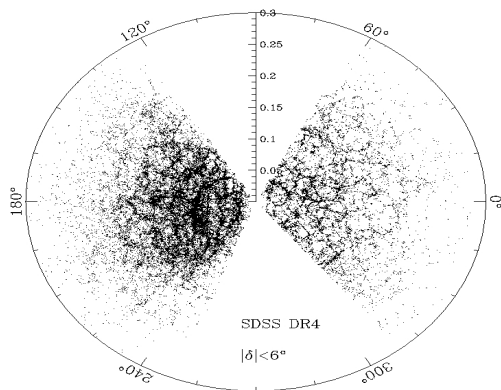
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# LSS of Universe

- Structures of the scale of Mpc and higher. 1 Parsec  
=  $3 * 10^{18}$ cm
- Formed due to the presence of small inhomogeneities in the early Universe
- Evolved primarily due to gravitational interaction among matter particles

# LSS Distribution



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- Comparison of results in galaxy catalog with model predictions

# Aims of Statistical Analysis

- \* Cosmographical Description of LSS.
- \* Physics of structure formation
  - Direct relation with nature of primordial fluctuation
  - Knowledge about DM distribution with the help of bias.
- \* The scale of crossover to homogeneity
- \* Form of galaxy correlation on small scales
- \* Complete Statistical Information about the distribution

## Two Point Correlation function

In a homogeneous distribution

$$dP = \frac{\bar{n}}{N} dV,$$

Hence

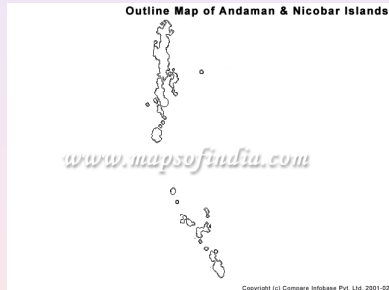
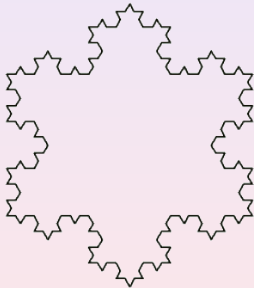
$$dP_{12} \propto dP_1 dP_2,$$

In a clustered distribution

$$dP_{12} = \left( \frac{\bar{n}}{N} \right)^2 [1 + \xi(\vec{r}_1, \vec{r}_2)] dV_1 dV_2$$

# Fractals

Fractal is an object made of parts similar to the whole structure in some sense.



These are irregular objects defying the geometric measures.



# Fractal Dimension

- For strictly self similar objects, Dimension  $D$  is

$$N = S^{-D}$$

$N$  = No of miniatures ;  $S$  = Scale factor Fractals are the objects which have fractional dimension.

- For strictly not self similar objects

$$N(r) \sim r^{-D}$$

$N(r)$  : No. of hyper-cubes required to do covering

# Correlation Dimension and Generalisation

Correlation integral :

$$C_2(r) = \frac{1}{N^2} \sum_{i=1}^N n_i(r)$$

The Dimension  $D_2$ , then is :

$$C_2(r) \sim r^{D_2}$$

Generalising the relation we get :

$$C_q(r) = \frac{1}{NM} \sum_{i=1}^M n_i^{q-1}(r)$$

The generalised Dimension  $D_q$ , then is :

$$C_q(r) \sim r^{D_q(q-1)}$$

# Minkowski-Bouligand Dimension

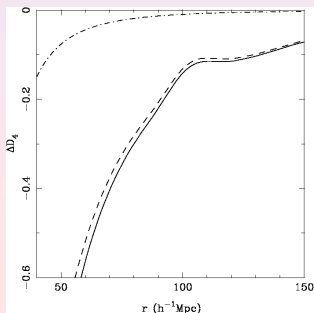
$$D_q = \frac{1}{q-1} \frac{\partial \log C_q}{\partial \log r}$$

- 1 Different in different range of scales
- 2  $q = 1, 2$  are Box counting dimension the correlation dimension
- 3 Complete statistical information by way of all the higher correlation function
- 4  $D_q = D$  for a monofractal
- 5 Information about regions with various amount of clustering

# Minkowski-Bouligand Dimension for a Weakly-Clustered Distribution

$$D_q(r) = D - \frac{D(q-2)}{\bar{N}} - \frac{Dq}{2} (\bar{\xi}(r) - \xi(r))$$

$$\Delta D_q = -(\Delta D_q)_{\bar{N}} - (\Delta D_q)_{clus}$$



With

$$b = 2, \bar{N} = 5 \times 10^{-5}$$

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(2008), MNRAS, 390,  
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Dispersion in  $\Delta D_q$ 

$$\text{Var}\{Dq\} \simeq \text{Var}\{(\Delta D_q)_{clus}\}$$

This implies that

$$\begin{aligned} \text{Var}\{\Delta D_q\} &\simeq \text{Var}\left\{\frac{Dq}{2} (\bar{\xi}(r) - \xi(r))\right\} \\ &\simeq \left(\frac{Dq}{2}\right)^2 (\text{Var}\{\bar{\xi}(r)\} + \text{Var}\{\xi(r)\} + 2\text{Cov}\{\bar{\xi}(r)\xi(r)\}) \\ &\simeq \left(\frac{Dq}{2}\right)^2 (\text{Var}\{\bar{\xi}(r)\} + \text{Var}\{\xi(r)\} \\ &\quad + 2\sqrt{\text{Var}\{\bar{\xi}(r)\}\text{Var}\{\xi(r)\}}) \\ &\simeq \left(\frac{Dq}{2}\right)^2 \left(\sqrt{\text{Var}\{\bar{\xi}(r)\}} + \sqrt{\text{Var}\{\xi(r)\}}\right)^2 \end{aligned}$$

# Variance of Power Spectrum and Correlation Function

FKP (1994) proposed

$$\sigma_P(k) = \sqrt{\frac{2}{V}} \left( P(k) + \frac{1}{\bar{n}} \right),$$

With This

$$\text{Cov}_\xi(r, r') = \int \frac{dk k^2}{2\pi^2} j_0(kr) j_0(kr') \sigma_P^2(k)$$

We can relate this to

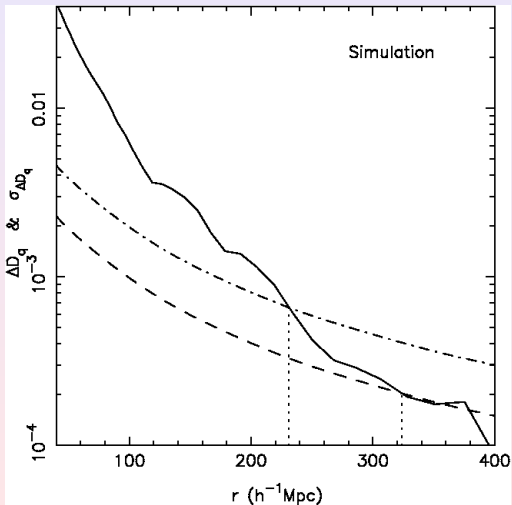
$$\begin{aligned} \text{Cov}_{\bar{\xi}}(i, j) &= \frac{1}{V_i V_j} \int d^3 r \int d^3 r' \text{Cov}_\xi(r, r') \\ &= \int \frac{dk k^2}{2\pi^2} \bar{j}_0(k, i) \bar{j}_0(k, j) \sigma_P^2(k) \end{aligned}$$

## Model Predictions

The *scale of homogeneity* can be defined as the scale above which  $\sigma_{\Delta D_q}$  exceeds  $\Delta D_q$

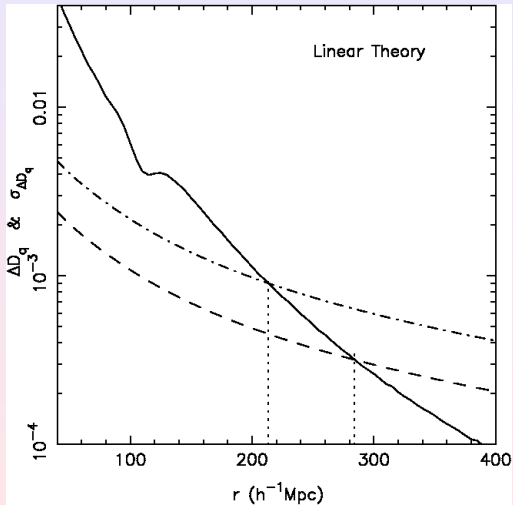
- As long as non-linear correction are not important, the scale of homogeneity does not change with epoch.
- In real space, the scale of homogeneity is independent of the tracer used as the deviation  $\Delta D_q$  as well as the dispersion in this quantity scale in the same manner with bias.
- Redshift space distortions introduce some bias dependance in the scale of homogeneity.
- As long as our assumption of  $q\bar{\xi} \ll 1$  is valid, the scale of homogeneity is the same for all  $q$ .

# N-Body Simulations





# Linear Theory



# Simulation + Missing Modes

